

# The energy flux associated with solar radiation incident

Question: Problem 12.014 The energy flux associated with solar radiation incident on the outer surface of the earth's atmosphere has been accurately measured and is known to be  $1368 \text{ W/m}^2$ . The diameters of the sun and earth are  $1.39 \times 10^9$  and  $1.27 \times 10^7$  m, respectively, and the distance between the sun and the earth is  $1.5 \times 10^{11}$  m (a) What is ...

Using the mean earth-sun distance, we can find the average value of solar irradiance incident to the earth's atmosphere, which is essentially constant. This value is therefore referred to as the solar constant, or air mass zero (AM0) ...

GLOBAL LONGWAVE RADIATION AT EDMONTON, ALBERTA. Gerald W. Sadler, in Energy Developments: New Forms, Renewables, Conservation, 1984 INTRODUCTION. The thermal radiant energy flux incident upon the surface of the earth may be conveniently separated into two components. The solar flux that arrives within the 0.3 to 3.5 micrometer wavelength range and ...

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A. Solar Radiation 1. Solar Constant 2. Spectral Composition of Sunlight a. Planck's Law b. Wien's Law c. absorption, reflection and transmission d. uv, PAR, NIR, IR L5.1 Introduction The sun is the source of energy that drives the cycle of life and death on earth. It is also the energy source that gives us warmth and evaporates water and ...

Earth intercepting the solar energy flux is  $\pi a^2$  where  $a$  is the radius of the Earth (Fig. 2.5), Solar power incident on the Earth  $= S_0 \pi a^2 = 1.74 \times 10^{17} \text{ W}$  using the data in Table 1.1. Not all of this radiation is absorbed by the Earth; a significant fraction is reflected. The ratio of reflected to incident solar energy is called the albedo,  $\alpha$ .

Radiation emitted by the sun is referred to as solar or shortwave radiation. Shortwave radiation - refers to the wavelength band ( $< 4 \mu\text{m}$ ) that carries most of the energy associated with solar radiation Solar constant (or total solar irradiance) ( $S_0$ ): The solar radiative flux, perpendicular to the solar beam, that enters

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the top of the ...

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(4) The energy flux associated with solar radiation incident on the outer surface of the earth's atmosphere has been accurately measured and is known to be  $1,400 \text{ W/m}^2$ . The diameters of the sun and earth are  $1.39 \times 10^9$  and  $1.29 \times 10^7$  m, respectively, and the distance between the sun and the earth is  $1.5 \times 10^{11}$  m (a) What is the emissivity power of the sun?

1) The energy flux associated with solar radiation incident on the outer surface of the earth's atmosphere has been accurately measured and is known to be  $1368 \text{ W/m}^2$ . The diameters of the sun and earth are  $1.39 \times 10^9$  m and  $1.27 \times 10^7$  m, respectively, and the distance between the sun and the earth is  $1.5 \times 10^{11}$  m. (30 points) a.

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We know that the energy received by the earth is equal to the energy emitted by the sun. The distance between the sun and the earth is  $1.5 \times 10^{11}$  m, and the energy flux associated with solar radiation incident on the outer surface of the earth's atmosphere is  $1366 \text{ W/m}^2$ . Therefore, the power received by the earth is:

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**Solar Flux Density Reaching Earth Solar Constant (S)** The solar energy density at the mean distance of Earth from the sun ( $1.5 \times 10^{11}$  m)  $S = L / (4 \pi d^2) = (3.9 \times 10^{26} \text{ W}) / [4 \times 3.14 \times (1.5 \times 10^{11} \text{ m})^2] = 1370 \text{ W/m}^2$   
ESS55 Prof. Jin-Yi Yu Solar Energy Incident On the Earth Solar energy incident on the Earth = total amount of solar energy can be ...

In other words, the farther away an object is from the sun, the lower the power incident to the object's surface (as discussed in Photon Energy and Flux). Using the mean earth-sun distance, we can find the average value of solar irradiance incident to the earth's atmosphere, which is essentially constant.

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2) The energy flux associated with solar radiation incident on the outer surface of the earth's atmosphere has been accurately measured and is known to be  $1368 \text{ W/m}^2$ . The diameters of the sun and earth are  $1.39 \times 10^9$  and  $1.27 \times 10^7$  m, respectively, and the distance between the centers of the sun and the earth is  $1.5 \times 10^{11}$  m.

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